Covariant Symplectic Structure and Conserved Charges of NMG

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Basics of Symplectic Geometry

Application to a Local Gravity Action

Calculation of Conserved Charges

Conclusion

Definition of Symplectic Structure

 Consider a smooth manifold Z endowed with a 2-form given as

$$\omega = d\mathbf{p}_i \wedge d\mathbf{q}^i, \tag{1}$$

where q^i , p_i are coordinates and momenta, i = 1, ..., N.

$$\omega = \frac{1}{2} \omega_{IJ} dQ^I \wedge dQ^J, \qquad (2)$$

with I = 1, ..., 2N. $Q^i = p_i$ for $i \le N$ and $Q^i = q^{i-N}$ for $i \ge N$. Then, $\omega_{i,i+N} = -\omega_{i+N,i} = -1$

Poisson bracket of any two function A(Q) and B(Q) is given by

$$[\mathbf{A},\mathbf{B}] = \omega^{IJ} \frac{\partial \mathbf{A}}{\partial Q^{I}} \frac{\partial \mathbf{B}}{\partial Q^{J}}.$$
 (3)

- ω is closed ($d\omega = 0$)
- nondegenerate, i.e. when ω is written as a 2N × 2N matrix, it has an inverse.
- ► This closed 2-form on *Z* is called symplectic structure.

Symplectic Structure in Geometrical Theories

- Choosing p_i, qⁱ as coordinates of the phase space Z would destroy the general covariance (by the choice of time coordinate).
- One should construct the phase space Z from solutions of the equations of motion derived from an action to achieve a manifestly covariant structure.
- Since classical solutions of any physical theory is in one-to-one correspondence with the initial values of p_i, qⁱ, we define our phase space as the space of solutions of the classical equations. (Crnkovic,Witten-1987)

- Let g be a solution of the field equations Φ .
- The functions on Z, denoted by g(x), takes a spacetime point x and maps it into a D×D real matrix g(x).

- Consider an arbitrary, small variation in the metric $ilde{g} = g + \delta g$
- When inserted into the field equations, it yields $\tilde{\Phi} = \Phi + \delta \Phi$.
- The vectors can be defined as the variations δg which solve $\delta \Phi = 0$ (preserves the field equations).

Fundamental Objects on Z - Forms

- A one-form, $\delta g(x)$, maps a vector δg to a $D \times D$ real matrix $\delta g(x)$, which is the vector evaluated at a spacetime *x*.
- We generalize this notion to construct p-forms as "wedge functions" of the one forms δg(x)

$$\Omega = \int dx_1 \cdots dx_p \,\Theta(x_1, \cdots, x_p) \,\delta g(x_1) \wedge \cdots \wedge \delta g(x_p), \quad (4)$$

where $\Theta(x_1, \dots, x_p)$ is a zero-form on *Z* and \wedge is an anticommuting product.

We can define an exterior derivative operator δ which maps p-forms to p+1-forms as follows

$$\delta\Omega = \int dx_0 \, dx_1 \cdots dx_p \, \frac{\delta\Theta(x_1, \cdots, x_p)}{\delta g(x_0)} \, \delta g(x_0) \wedge \delta g(x_1) \wedge \cdots \wedge \delta g(x_p).$$
(5)

 One can easily check this operator obeys the modified Leibniz rule and Poincaré Lemma. Let us consider a generic local gravity action

$$S = \int d^{D}x \,\sqrt{|g|} \,\mathcal{L}(g, R, \nabla R, R^{2}, \cdots), \qquad (6)$$

whose variation is given as

$$\delta S = \int d^{D}x \,\sqrt{|g|} \,\Phi_{ab} \delta g^{ab} + \int d^{D}x \,\partial_{a} \Lambda^{a}(g, \delta g, \nabla \delta g \cdots), \quad (7)$$

where Φ_{ab} is the field equation and Λ^a is the boundary term.

- We can view δS as a 1-form on Z (note that $\Lambda^a(x)$ includes δg_{ab} and the relevant quantities).
- The exterior derivative of (7) will vanish by Poincaré Lemma,

$$\delta^{2}S = \int d^{D}x \,\sqrt{|g|} \,\delta\Phi_{ab} \wedge \delta g^{ab} - \frac{1}{2} \int d^{D}x \,\sqrt{|g|} \,\Phi_{ab} \,\delta g^{ab} \wedge \delta \ln|g| + \int d^{D}x \,\partial_{a} \delta\Lambda^{a} = 0.$$
(8)

where $\delta \ln|g| = g^{ab} \delta g_{ab} = -g_{ab} \delta g^{ab}$.

First two integrals vanish on shell and the third one implies

$$\int d^D x \,\sqrt{|g|} \,\nabla_a J^a = 0, \qquad (9)$$

where $J^a \equiv -\frac{\delta \Lambda^a}{\sqrt{|g|}}$ is the "symplectic current".

From this, one can construct the following Poincaré invariant 2-form since the covariant divergence of the symplectic current vanishes (∇_aJ^a = 0)

$$\omega = \int_{\Sigma} d\Sigma_a \sqrt{|g|} J^a , \qquad (10)$$

where Σ is (D-1)-dimensional spacelike hypersurface.

Darboux's theorem assures us that this is the sought after symplectic structure of the theory if ω is additionally closed.

- The former is trivial since all constituents of ω transform like tensors. For the latter, we should find out how ω transforms under the following transformation

$$\delta g_{ab} \to \delta g_{ab} + \nabla_a \xi_b + \nabla_b \xi_a \tag{11}$$

where ξ is asymptotic to a Killing vector field at infinity. This computation will yield a boundary term which gives rise to conserved charges of the theory under consideration.

 Let us now apply this procedure to the following quadratic action

$$I = \int d^{D}x \ \sqrt{|g|} \mathcal{L} \equiv \int d^{D}x \ \sqrt{|g|} \left(\frac{1}{\kappa}(R + 2\Lambda_{0}) + \alpha R^{2} + \beta R_{ab}^{2}\right)$$
(12)

 One can explicitly show that covariant divergence of the symplectic current is equal to the following expression

$$\nabla_{a}J^{a} = \frac{1}{2}g^{ab}\delta\Phi_{ab}\wedge\delta\ln|g| + \delta\Phi_{ab}\wedge\delta g^{ab} + \delta\Phi\wedge\delta\ln|g|,$$
(13)

which vanishes on shell. Here $\Phi_{ab} \equiv \frac{1}{\kappa} \mathcal{G}_{ab} + \alpha A_{ab} + \beta B_{ab}$ and $\Phi = g^{ab} \Phi_{ab}$.

It can also be shown, without using the field equations, that ω is a closed form. We have

$$\delta\omega = \int_{\Sigma} d\Sigma_a \, (\delta \sqrt{|g|} \wedge J^a + \sqrt{|g|} \, \delta J^a), \tag{14}$$

and variation of the current reads

$$\delta J^{a} = -\frac{1}{2}J^{a} \wedge \delta \ln|g|. \tag{15}$$

By virtue of (15) and bearing in mind that J^a is anticommuting 2-form, (14) vanishes.

There remains to investigate the gauge invariance of ω. After a cumbersome calculation, change in the symplectic current can be written as

$$\Delta J^{a} = \nabla_{c} \mathcal{F}^{ac} + g^{bc} \delta \Phi_{bc} \wedge \xi^{a} + 2\Phi_{bc} \xi^{c} \wedge \delta g^{ab} + \Phi^{ac} \xi_{c} \wedge \delta \ln|g| + \xi^{a} \wedge \delta \Phi$$
(16)

First term in (16) vanishes when inserted in the integral for ω for sufficiently fast decaying metric variations, the remaining terms vanish on-shell.

where $\mathcal{F}^{ac} = -\mathcal{F}^{ca} = \frac{1}{\kappa} \mathcal{F}^{ac}_{\kappa} + \alpha \mathcal{F}^{ac}_{\alpha} + \beta \mathcal{F}^{ac}_{\beta}, \qquad (17)$

with

$$\mathcal{F}_{\kappa}^{ac} \equiv 2\xi^{[c} \wedge \nabla_{b}\delta g^{a]b} - 2\xi_{b} \wedge \nabla^{[c}\delta g^{a]b} - 2\delta g^{b[c} \wedge \nabla_{b}\xi^{a]} - 2\xi^{[a} \wedge \nabla^{c]}\delta \ln|g| - \delta \ln|g| \wedge \nabla^{[c}\xi^{a]}$$
(18)

Calculation of Conserved Charges

- We linearize the metric as $g_{ab} = \bar{g}_{ab} + h_{ab}$
- Indices are raised/lowered and covariant derivatives are defined with respect to the background metric g
 _{ab} as usual.
- One should take the diffeomorphisms as the isometries of the background spacetime, meaning $\overline{\nabla}_a \overline{\xi}_b + \overline{\nabla}_b \overline{\xi}_a = 0$.
- Assume the background spacetime g
 _{ab} admits a globally defined Killing vector ξ
 _a.
- The variation is identified as δg_{ab} → h_{ab}, δg^{ab} → −h^{ab}. Therefore, the terms R_{ab}, R are identified with the background ones R

 _{ab}, R

 and terms like δ(∇_aR_{bc}) are taken as (∇_aR_{bc})_L, where subscript L means linearized version of the corresponding quantity.
- Finally, we write the ξ terms at the right hand side of the wedge products and then drop them.

Calculation of Conserved Charges

 With all these identifications the relevant charge expression is given by

$$Q(\bar{\xi}) = \frac{1}{2} \int_{\Sigma} d^{D-1} x \sqrt{|\sigma|} n_a \, \bar{\nabla}_c Q^{ac}$$

= $\frac{1}{2} \int_{\partial \Sigma} d^{D-2} x \sqrt{|\sigma^{(\partial \Sigma)}|} n_a \, s_c \, Q^{ac}, \qquad (19)$

where Σ is a (D-1)-dimensional spacelike hypersurface with induced metric σ and unit normal vector n^a , $\partial \Sigma$ (boundary of Σ) is a (D-2)-dimensional hypersurface with induced metric $\sigma^{(\partial \Sigma)}$ and unit normal s^c .

Calculation of Conserved Charges

$$Q^{ac} = -Q^{ca} = \frac{1}{\kappa} Q^{ac}_{\kappa} + \alpha Q^{ac}_{\alpha} + \beta Q^{ac}_{\beta}, \qquad (20)$$

with

$$Q_{\kappa}^{ac} \equiv 2\bar{\nabla}_{b}h^{b[a}\bar{\xi}^{c]} - 2\bar{\nabla}^{[c}h^{a]b}\bar{\xi}_{b} - 2h^{b[c}\bar{\nabla}_{b}\bar{\xi}^{a]} + 2(\bar{\nabla}^{[c}h)\bar{\xi}^{a]} - h\bar{\nabla}^{[c}\bar{\xi}^{a]}$$
(21)

$$\mathcal{F}_{\kappa}^{ac} \equiv 2\xi^{[c} \wedge \nabla_{b}\delta g^{a]b} - 2\xi_{b} \wedge \nabla^{[c}\delta g^{a]b} - 2\delta g^{b[c} \wedge \nabla_{b}\xi^{a]b} - 2\xi^{[a} \wedge \nabla^{c]}\delta \ln|g| - \delta \ln|g| \wedge \nabla^{[c}\xi^{a]}$$

(22)

- Charge expressions given above are identical to the ones given in the literature.
- Remember that we assumed δφ_{ab} = 0 while constructing the vectors. Here, this condition implies (Φ_{ab})_L = 0 at the boundary of the spacetime.

Calculation of Conserved Charges - Examples

3-dimensional Lifshitz Blackhole

$$ds^{2} = -\frac{r^{6}}{\ell^{6}} \left(1 - \frac{M\ell^{2}}{r^{2}}\right) dt^{2} + \frac{\ell^{2}}{r^{2}} \left(1 - \frac{M\ell^{2}}{r^{2}}\right)^{-1} dr^{2} + \frac{r^{2}}{\ell^{2}} dx^{2}, \quad (23)$$

$$\Lambda_0 = \frac{13}{2\ell^2}, \quad \beta = \frac{2\ell^2}{\kappa}, \quad \alpha = -\frac{3\ell^2}{4\kappa}, \quad \kappa = 16\pi G$$

with background

$$ds^{2} = -rac{r^{6}}{\ell^{6}}dt^{2} + rac{\ell^{2}}{r^{2}}dr^{2} + rac{r^{2}}{\ell^{2}}dx^{2},$$

Calculation of Conserved Charges - An Example

$$n^{a} = -\frac{\ell^{3}}{r^{3}}\delta^{a}_{t}, \ s^{a} = \frac{r}{l}\delta^{a}_{r}$$
$$E = \lim_{r \to \infty} \int_{0}^{2\pi\ell} \frac{r^{3}}{\ell^{3}} n_{t} s_{r} Q^{tr}(\bar{\xi}) dx = \frac{7m^{2}}{4G}, \qquad (24)$$
$$(\Phi_{ab})_{L} = \frac{6M^{3}}{\ell^{2}} \qquad (25)$$

at the boundary.

- We have constructed the symplectic structure of NMG and calculated the conserved charges of some solutions using the diffeomorphism invariance of the symplectic two-form.
- This method gives rise to the condition (Φ_{ab})_L = 0 at the boundary, which the solutions with AdS background satisfy.
- However, solutions with arbitrary background which we have considered do not satisfy the condition (Φ_{ab})_L = 0 at the boundary. This might be the reason for the discrepancy in the thermodynamics of these blackholes. (Devecioglu,Sarioglu-2011)